W3OH spectrum from 210 to 272GHz with IRAM-30m

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Abstract

The calibration source CALW3OH (Mauersberger et al. 1989, A&ASS 79, 217) was observed with the IRAM-30m facility and its EMIR band 2 (230GHz) receiver. We used the FTS spectrometer with 200 kHz resolution covering 2×8 GHz (USB and LSB) in each polarization with EMIR 2 which operates in DSB mode. The 1.3mm band was nearly fully covered with 5 different setups, with local oscillator set at 222300, 230080, 236600, 252080 and 260425 MHz. The source was observed as a calibrator for comet observations during the observing runs of program 115-12. Beam efficiency (η_{mb}) was checked mainly on Uranus and spectra have been corrected for additionnal losses due to variation of efficiency as a function of elevation. The spectra are given in Tmb scale. Observation were obtained on the 14th of March, 6th, 7th and 8th of April 2013. We have selected the observations with best focus (FWHM and Beam efficencies close to nominal, i.e. η_{mb} between 0.54 and 0.43 from 210 to 270 GHz) and elevation was between 32 and 64°. Atmospheric precipitable water vapour ranged from 1 to 6 mm. The observing mode used was position switching every 30 s with a reference at 600". Several spectra were averaged, baseline subtracted, and spectra with poor baselines (e.g. ripples in one polarization) were removed.

1 The spectra

All spectra have been combined in one single spectrum spaning 62 GHz and is in 12×6 sections of 860 GHz. Vertical scale is automatically set by class software according to the strongest line in the window. Noise may vary depending on integration time and sky transmission at the time of observations.

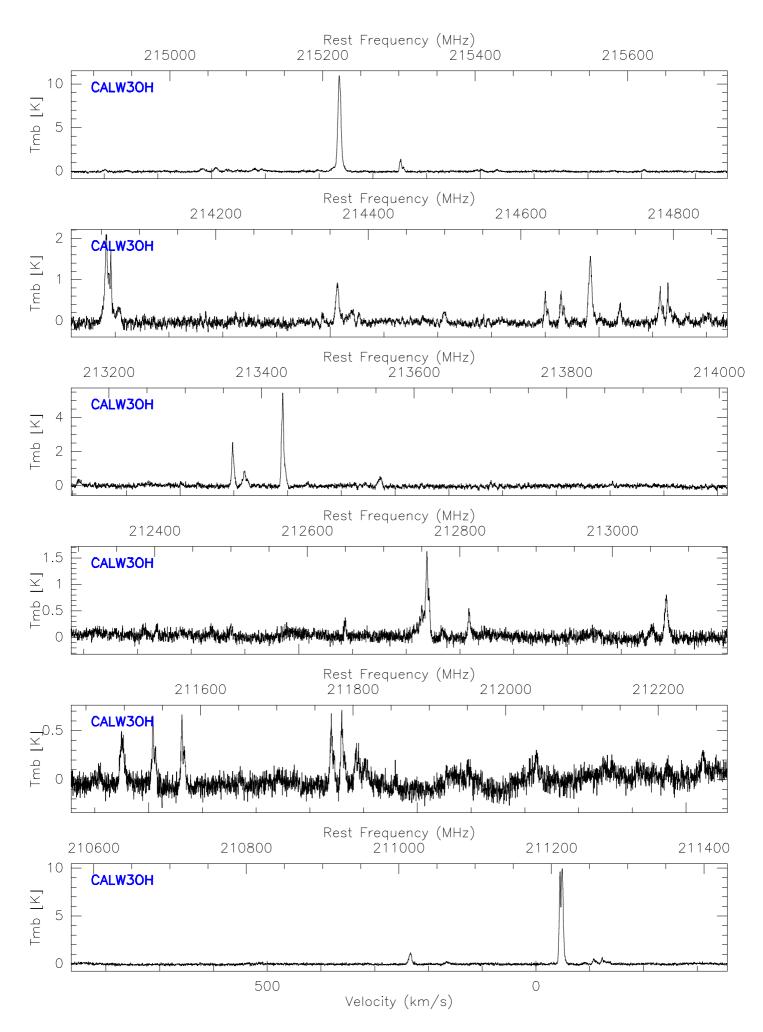


Figure 1:

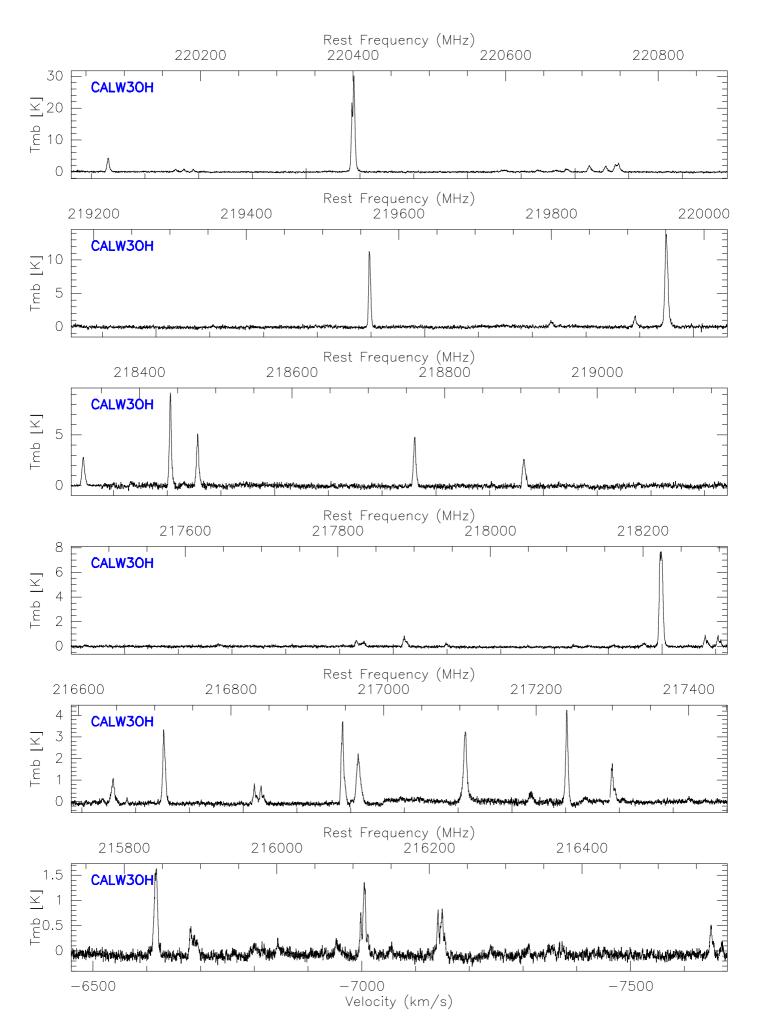


Figure 2:

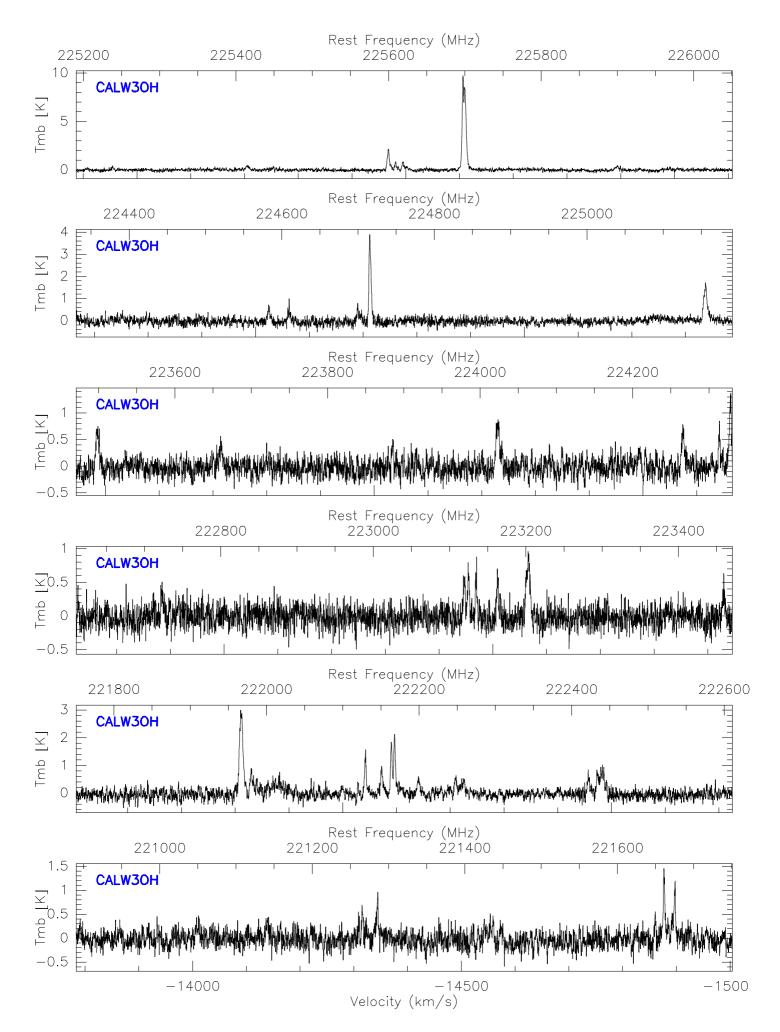


Figure 3:

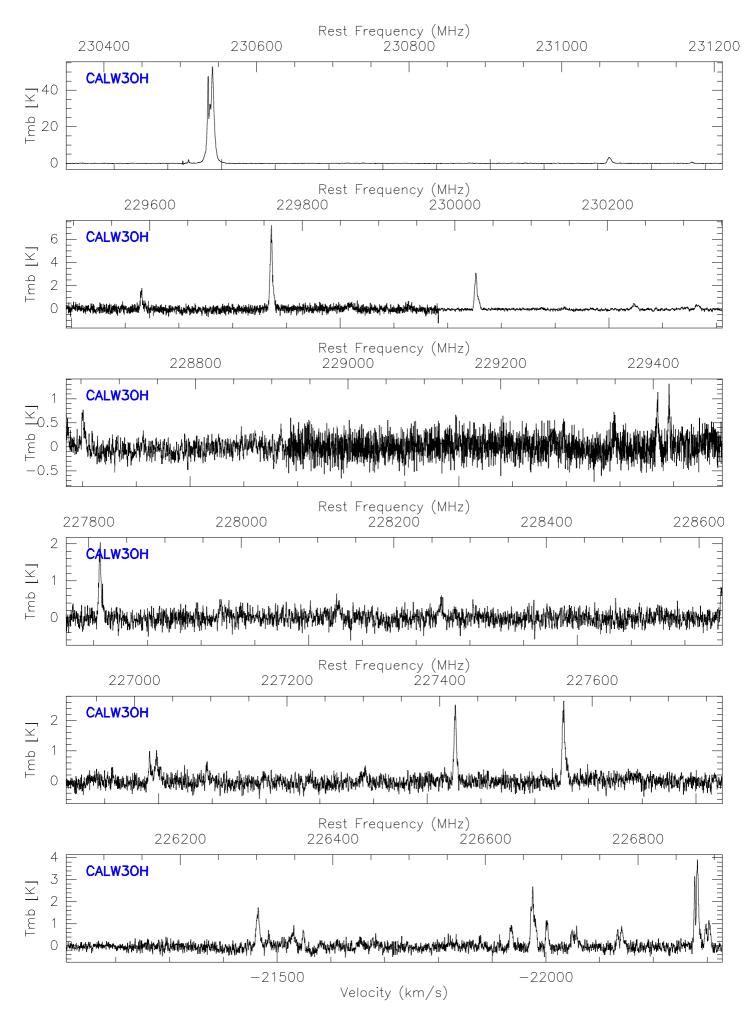


Figure 4:

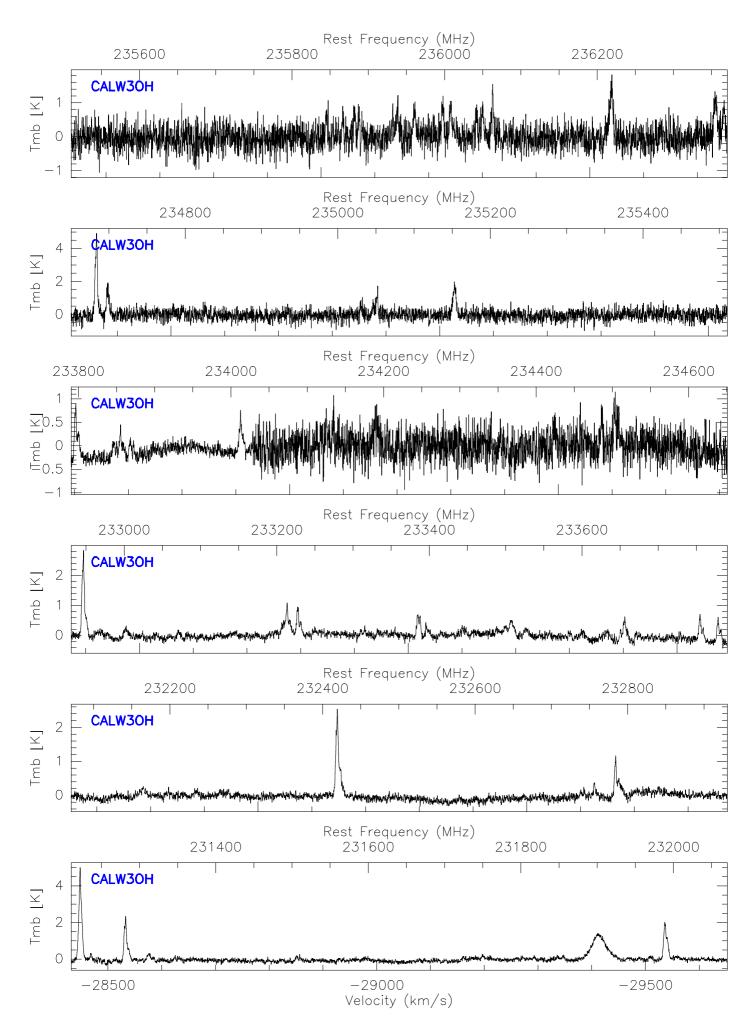


Figure 5:

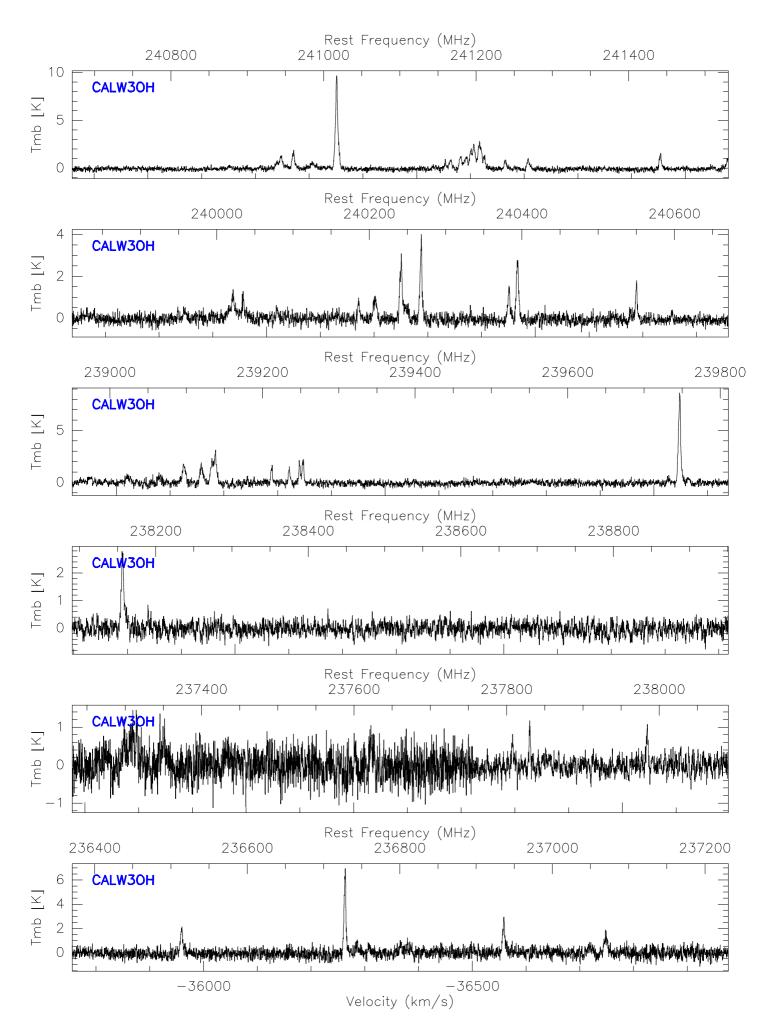


Figure 6:

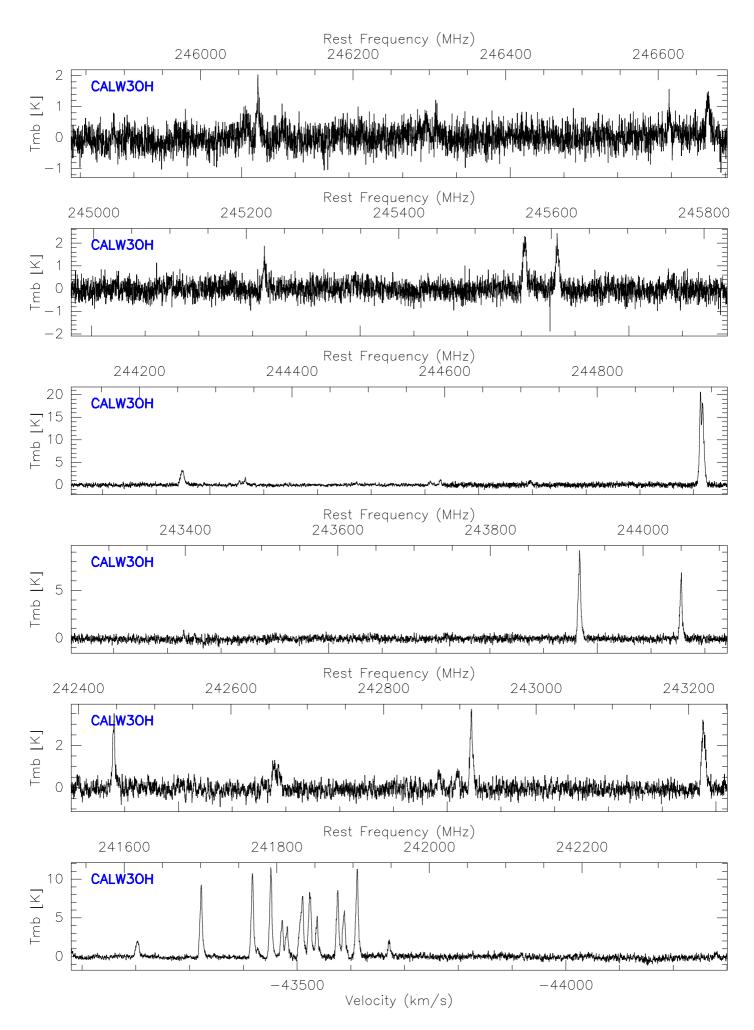


Figure 7:

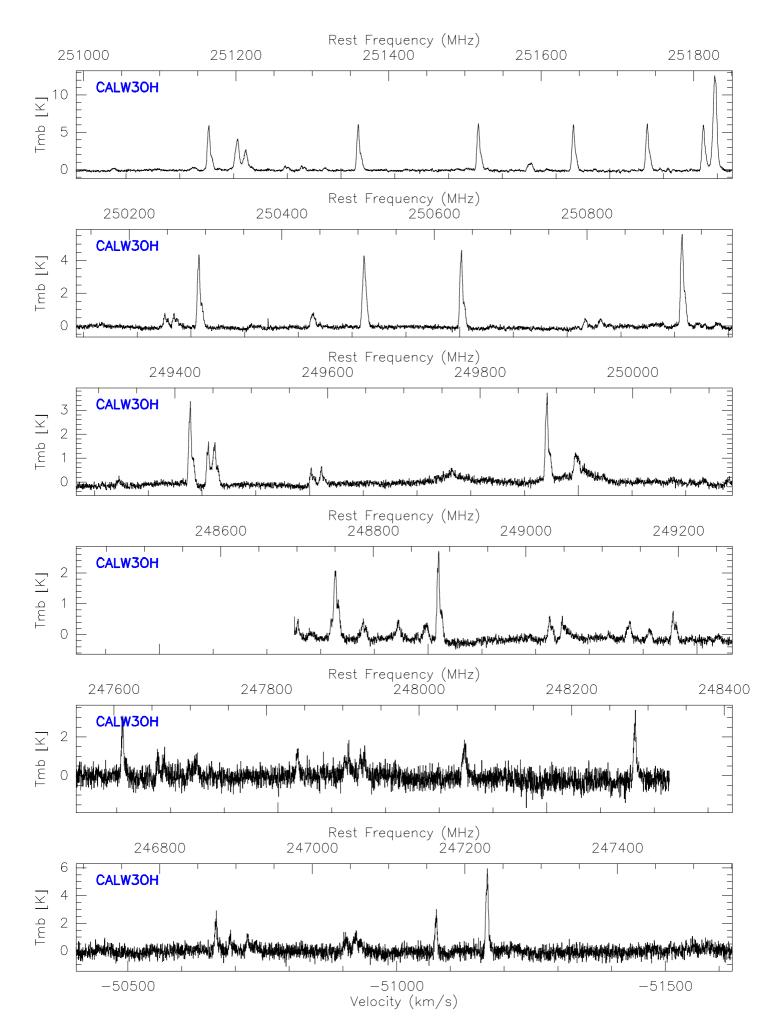


Figure 8:

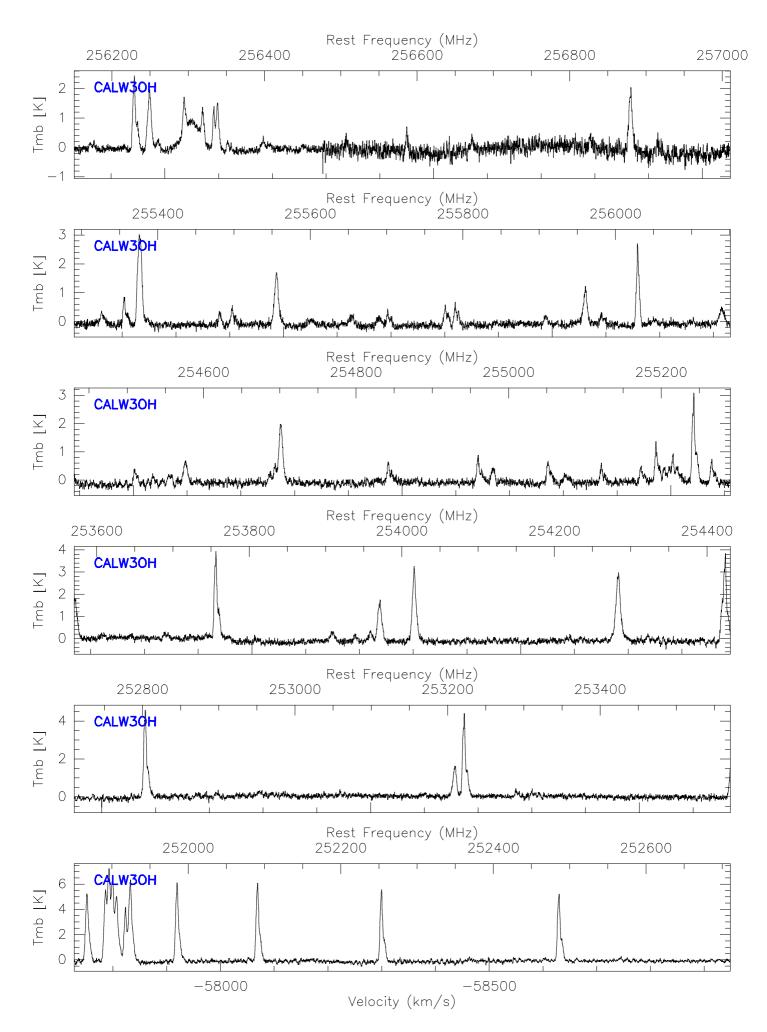


Figure 9:

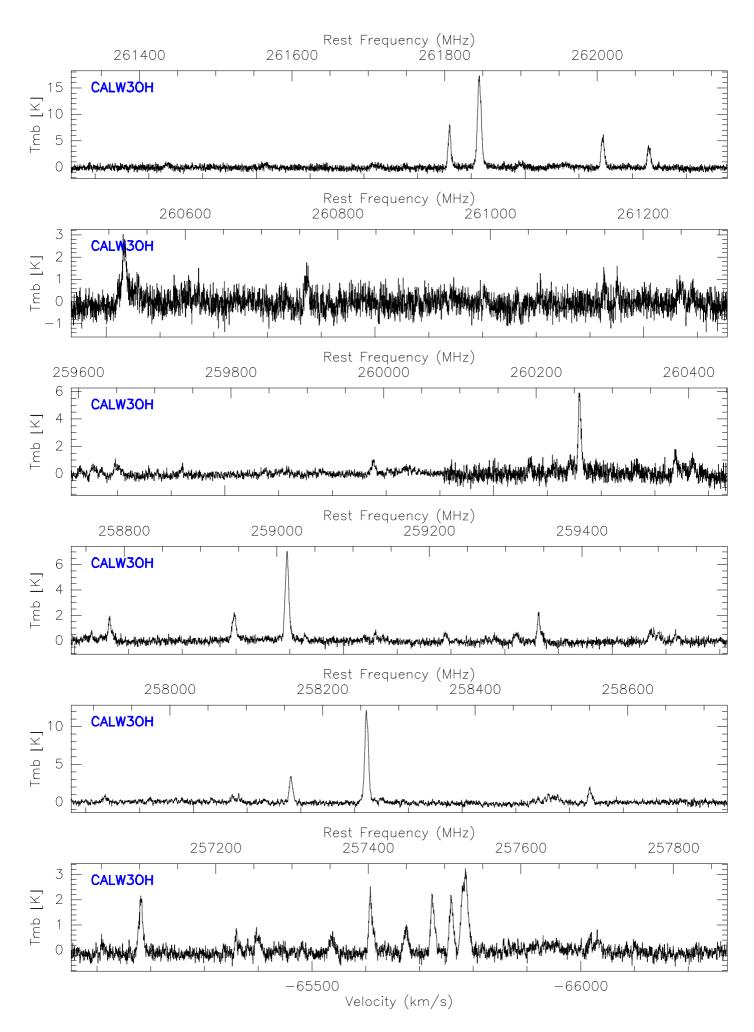


Figure 10:

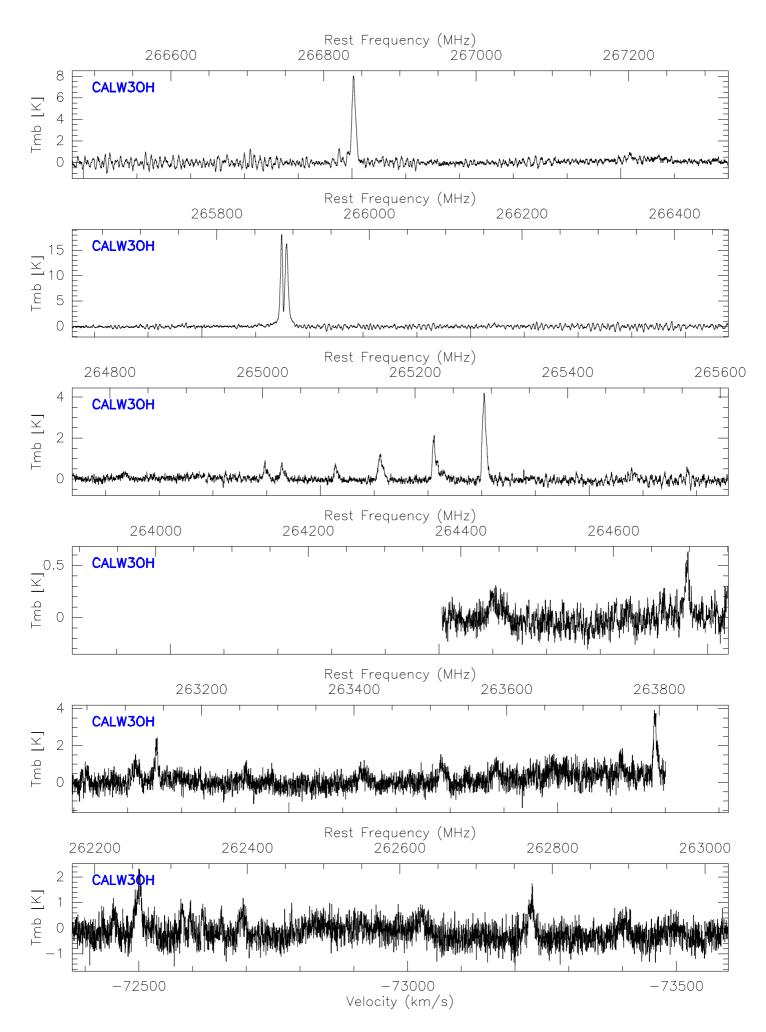


Figure 11:

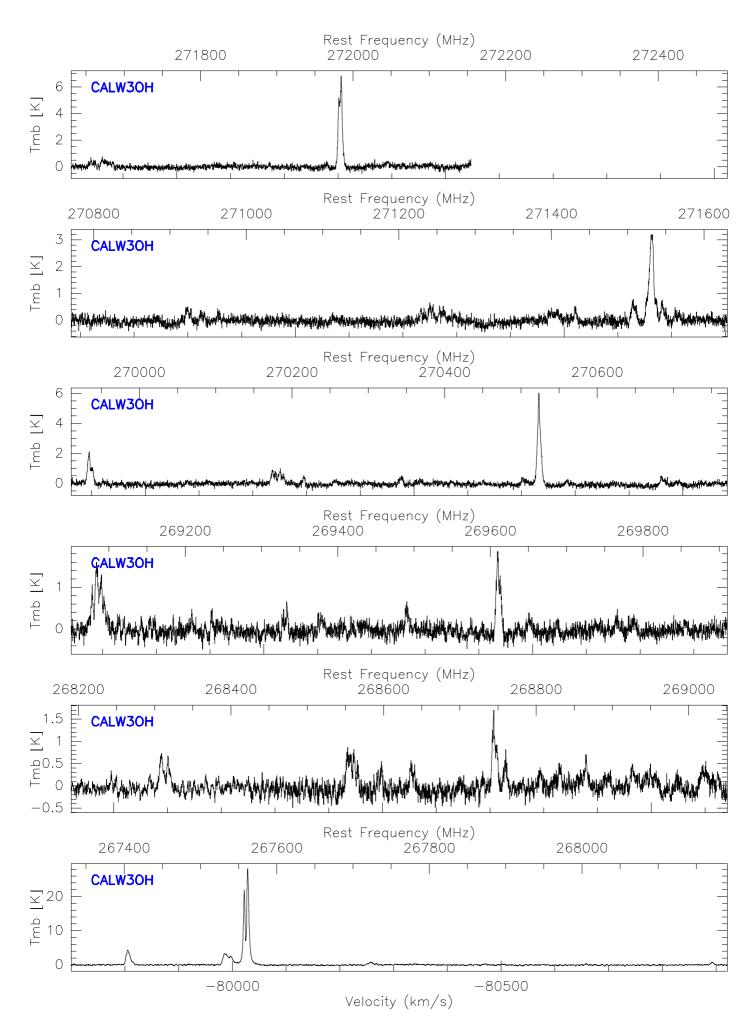


Figure 12: